

Dresden – March 5th, 2013

Discrete symmetries in supersymmetric models of flavor

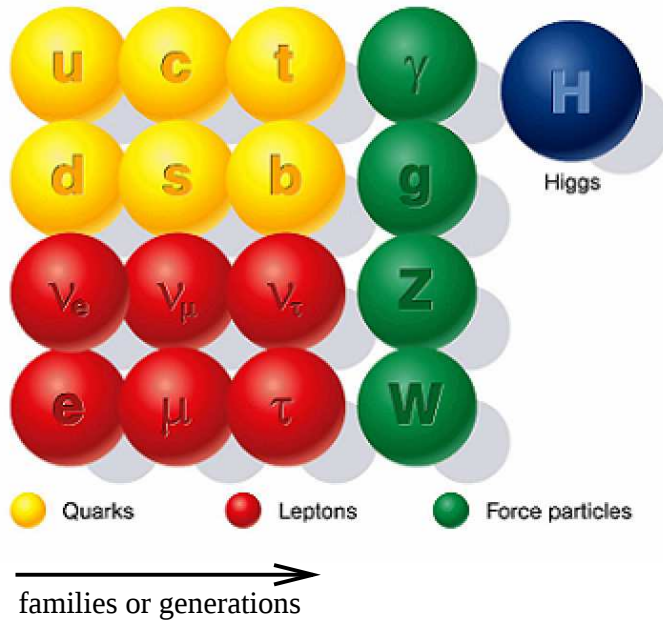
Christoph Luhn



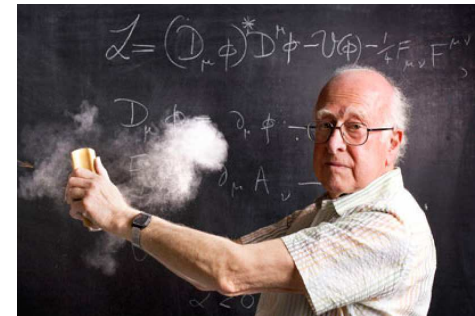
Outline

- ▶ supersymmetric Standard Model
 - Z_2 symmetries R -parity and matter parity
 - possible $U(1)_{B-L}$ origin
- ▶ Abelian family symmetries
 - $U(1)_{\text{FN}}$ for explaining hierarchies
 - residual discrete symmetries
 - R -symmetries
- ▶ non-Abelian family symmetries
 - motivation
 - promising candidates
 - direct and indirect implementation
 - benchmark example based on S_4

Standard Model (of particle physics)



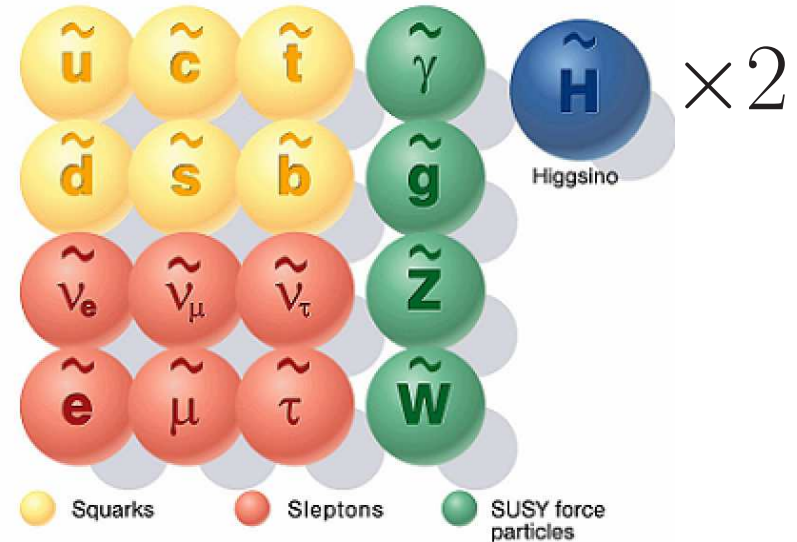
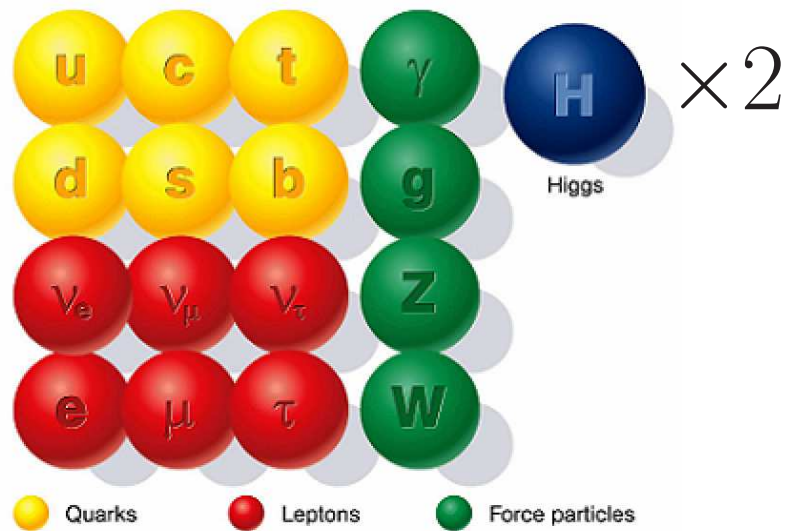
- highly successful theory
- based on gauge symmetry
 $SU(3)_C \times SU(2)_W \times U(1)_Y$
- broken by Higgs vacuum



Why look beyond ?

- hierarchy problem (protecting the Higgs mass)
- origin of three families of quarks & leptons
- neutrino masses and mixing
- baryogenesis, dark matter, dark energy, etc.

Supersymmetric Standard Model



- addresses hierarchy problem
- most general space-time symmetry (extension of Poincaré symmetry)
- supersymmetry must be broken – at TeV scale?
- in total > 300 parameters [Barbier et al. \(hep-ph/0406039\)](#)
- impose extra symmetries, e.g. R -parity $\sim$ matter parity
 $\rightarrow$ MSSM with 124 parameters [Haber \(hep-ph/9709450\)](#)

R -parity and matter parity

superfield formalism, e.g.

$$Q = \tilde{q} + q\theta + F_q\theta^2$$

q = quark

$\tilde{q}$ = squark

θ = superspace variable

- R -parity (R_p) defined on component fields, e.g. q and $\tilde{q}$

R_p	quarks	leptons	Higgs	squarks	sleptons	Higgsino
additive Z_2	0	0	0	1	1	1

- matter parity (M_p) defined on superfields, e.g. Q

M_p	(s)quarks	(s)leptons	Higgs(inos)
additive Z_2	1	1	0

R_p and M_P allow and forbid exactly the same terms in Lagrangian

Matter parity from $U(1)_{B-L}$

- matter parity
- forbids (renormalizable) B and L violation
 - introduced to stabilize proton
 - violated by quantum gravity effects [Krauss, Wilczek \(1989\)](#)
 - unless gauge origin, e.g. from breaking $U(1)_{B-L}$

	Q	U^c	D^c	L	E^c	H_u	H_d	ϕ
$U(1)_{B-L}$	$\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	-1	1	0	0	$\pm\frac{2}{3}$
$3 \times U(1)_{B-L}$	1	-1	-1	-3	3	0	0	± 2
M_p	1	1	1	1	1	0	0	0

→ vev of **field** ϕ breaks $U(1)_{B-L}$ spontaneously down to M_p

“Higgs”-type field

neutral under

SM gauge group

$U(1)_{\text{FN}}$ family symmetry Froggatt, Nielsen (1979)

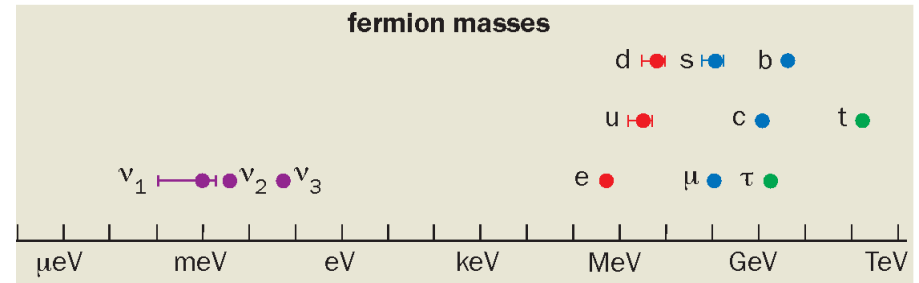
Fermion masses

- fermion mass terms: $M_{ij} \bar{\psi}_i \psi_j$
- masses = eigenvalues of M

$$m_u : m_c : m_t \sim \lambda^8 : \lambda^4 : 1$$

$$m_d : m_s : m_b \sim \lambda^4 : \lambda^2 : 1$$

$$m_e : m_\mu : m_\tau \sim \lambda^{4 \text{ or } 5} : \lambda^2 : 1 \quad (\lambda \sim 0.22)$$



- neutrinos massless in Standard Model
- observation of neutrino oscillation $\rightarrow m_\nu \neq 0$
- three scenarios:

$$m_{\nu_1} : m_{\nu_2} : m_{\nu_3} \sim \begin{cases} \lambda^x : \lambda : 1 & \text{(normal hierarchy)} \\ 1 : 1 : \lambda^x & \text{(inverted hierarchy)} \\ 1 : 1 : 1 & \text{(quasi degenerate)} \end{cases}$$

Froggatt-Nielsen mechanism

- Yukawa terms $\boxed{Y_{ij} \bar{\psi}_i \psi_j H}$ forbidden by $U(1)_{\text{FN}}$
- family dependent charges
- introduce **flavon field** ϕ which allows effective terms $\boxed{c_{ij} \bar{\psi}_i \psi_j H \left(\frac{\phi}{\Lambda}\right)^{x_{ij}}}$
- flavon ϕ acquires a VEV $\rightarrow Y_{ij} = c_{ij} \left(\frac{\langle\phi\rangle}{\Lambda}\right)^{x_{ij}}$

fields	Q_1	Q_2	Q_3	D_1^c	D_2^c	D_3^c	H_d	ϕ
$U(1)_{\text{FN}}$	6	4	0	5	3	3	-3	-2

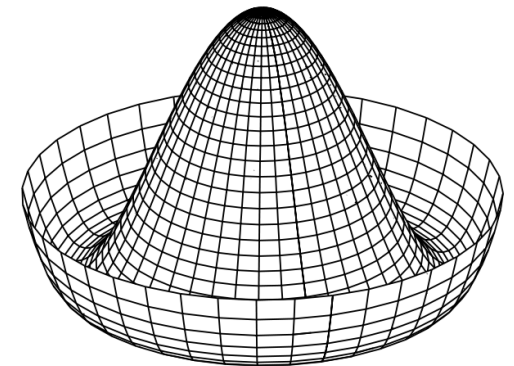
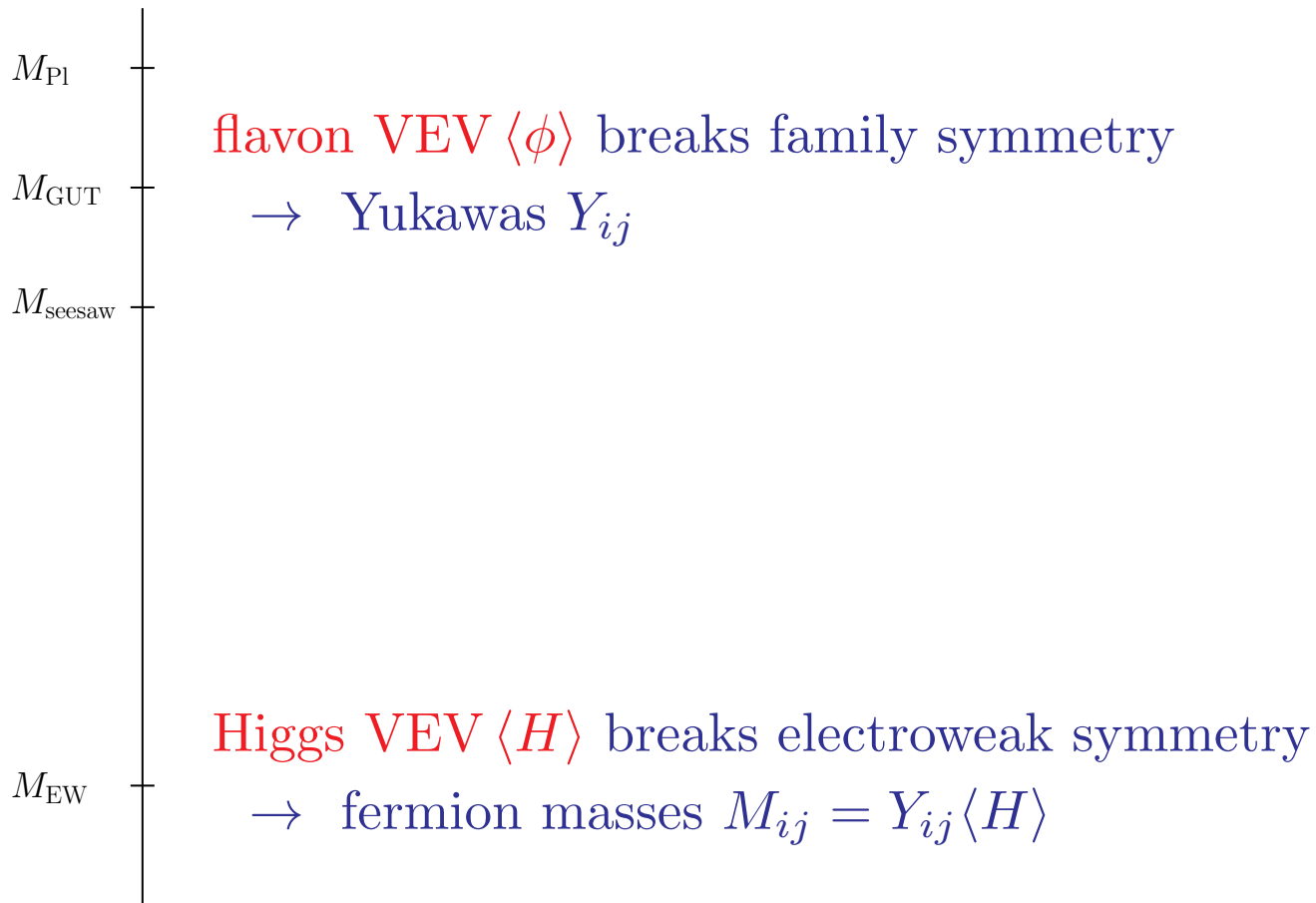
$$\frac{\langle\phi\rangle}{\Lambda} \sim \lambda$$

$$c_{ij} Q_i D_j^c H_d \left(\frac{\phi}{\Lambda}\right)^{x_{ij}} \rightarrow c_{ij} Q_i D_j^c H_d \lambda^{x_{ij}} \rightarrow Y_d \sim \begin{pmatrix} \lambda^4 & \lambda^3 & \lambda^3 \\ \lambda^3 & \lambda^2 & \lambda^2 \\ \lambda & 1 & 1 \end{pmatrix}$$

- hierarchies arise from spontaneous breakdown of $U(1)_{\text{FN}}$
- often residual discrete Z_N symmetry, e.g. matter parity
- $\mathcal{O}(1)$ coefficients c_{ij} not fixed

Scale of family symmetry breaking

- models of flavor are typically formulated at high energies
- separate family symmetry from EW symmetry breaking



Discrete symmetries from $U(1)_{\text{FN}}$

- simple conditions to obtain a particular Z_N symmetry from $U(1)_{\text{FN}}$
- in $U(1)_{\text{FN}}$ charge normalization with $X_\phi = -1$

	$X_{H_d} - X_{L_1}$	$3X_{Q_1} + X_{L_1}$
M_p	integer $-\frac{1}{2}$	integer
P_6	integer $-\frac{1}{2}$	integer $\pm \frac{1}{3}$
B_3	integer	integer $\pm \frac{1}{3}$

Dreiner, Luhn, Murayama,
Thormeier (hep-ph/0610026)

M_p lightest SUSY particle (LSP) is stable $\rightarrow$ dark matter candidate
allows for non-renormalizable operator $QQQL \rightarrow$ **proton decay**

P_6 LSP is stable $\rightarrow$ dark matter candidate
forbids $QQQL \rightarrow$ proton stabilized

B_3 LSP not stable $\rightarrow$ less missing E_T
forbids $QQQL \rightarrow$ proton stabilized
renormalizable L violation $\rightarrow$ neutrino masses without seesaw

Discrete Z_4^R symmetry from $U(1)_{\text{FN}}^R$

- superspace variable θ charged under R -symmetries

	Q	U^c	D^c	L	E^c	H_u	H_d	θ
Z_4^R	1	1	1	1	1	0	0	1

Lee et al.
(arXiv:1009.0905)

- solves the μ -problem by forbidding $\mu H_u H_d$
- forbids $QQQL$

-
- Z_4^R can be obtained from $U(1)_{\text{FN}}^R$ if

$X_{H_d} - X_{L_1}$	$3X_{Q_1} + X_{L_1}$	X_θ	X_ϕ
integer $\pm \frac{1}{4}$	integer	integer $\pm \frac{1}{4}$	-1

Dreiner, Luhn,
Opferkuch
(in preparation)

Non-Abelian family symmetries

Fermion mixings

- mismatch of flavor (weak) and mass eigenstates

$$\Psi_{\text{flavour}} = V^\dagger \Psi_{\text{mass}}$$

- quark sector: V_L^u and V_L^d

$$U_{\text{CKM}} = V_L^u V_L^{d\dagger} \approx \begin{pmatrix} 1 & \lambda & \lambda^3 \\ \lambda & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} \quad \lambda \sim 0.22$$

- lepton sector: V_L^e and V_L^ν

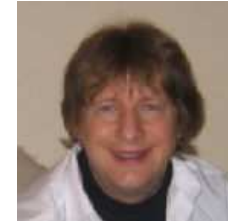
$$U_{\text{PMNS}} = V_L^e V_L^{\nu\dagger} \approx \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

Gonzalez-Garcia et al. (2012)

mixing $\iff$ each family knows of the existence of the others!

Tri-Bimaximal lepton mixing vs. global neutrino fits

$$U_{\text{PMNS}} \approx U_{\text{TB}} \equiv \begin{pmatrix} -\frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0 \\ \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}} \end{pmatrix}$$



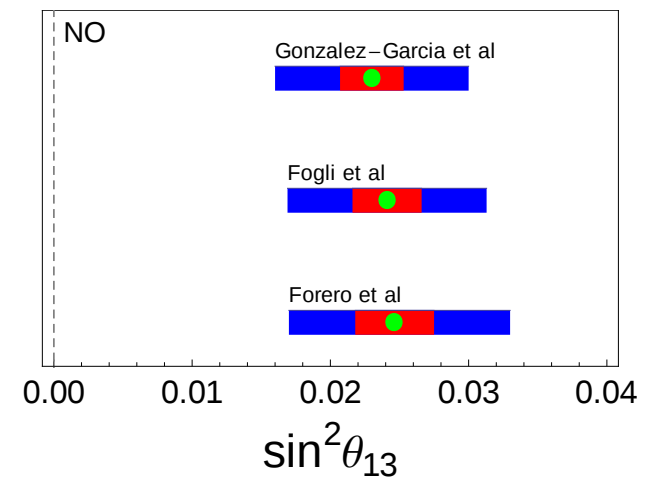
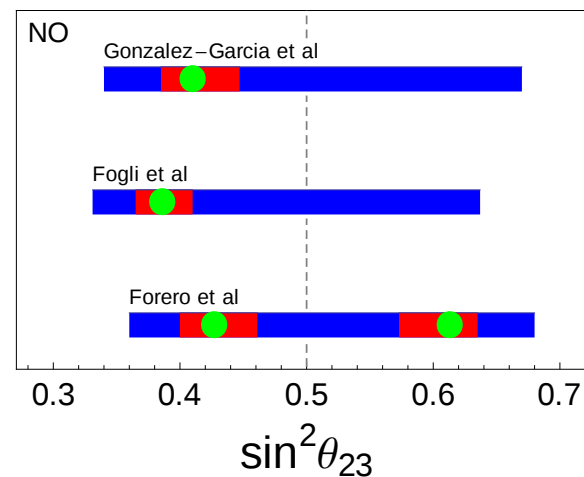
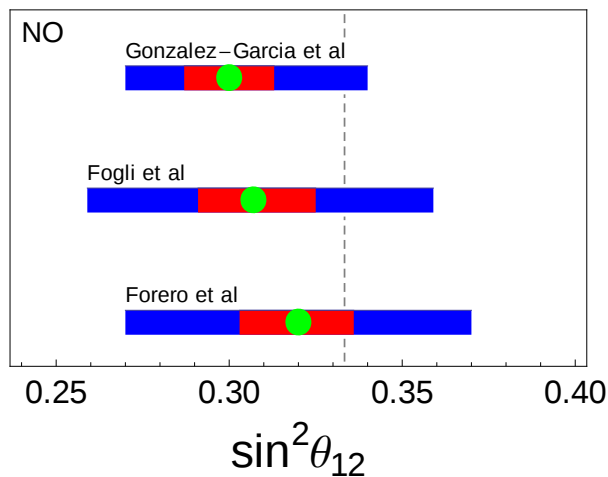
Harrison



Perkins



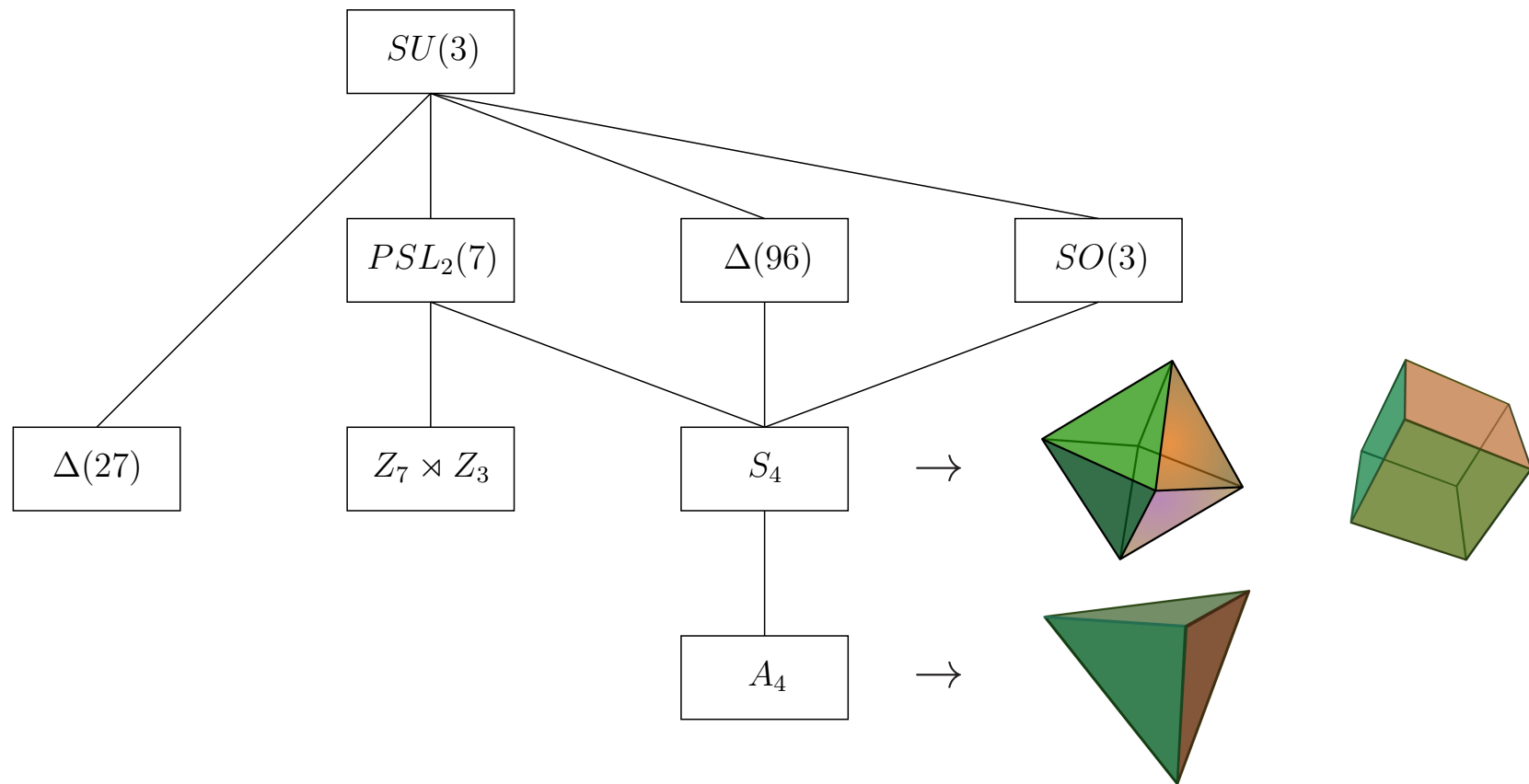
Scott



- TBmax mixing fits θ_{12} and θ_{23} very well – until recently also θ_{13}
- motivates family symmetries, e.g. A_4, S_4

Candidates

- unify three families in multiplets of family symmetry
- underlying group should have two- or three-dimensional representations



Symmetries of the mass matrices (in flavor basis)

charged leptons $M_\ell = \text{diag}(m_e, m_\mu, m_\tau)$



Dirac

symmetric under diagonal phase transformation h

$$\boxed{M_\ell = h^T M_\ell h^*} \quad \text{e.g. } h = \text{diag}\left(1, e^{\frac{4\pi i}{3}}, e^{\frac{2\pi i}{3}}\right)$$

neutrinos



Majorana

$$M_\nu = U_{\text{PMNS}} \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}) U_{\text{PMNS}}^T$$

symmetry of M_ν depends on U_{PMNS}

$$\boxed{M_\nu = k^T M_\nu k} \quad k = U_{\text{PMNS}}^* \text{diag}(+1, -1, -1) U_{\text{PMNS}}^T$$

four different $k \rightarrow$ generate $Z_2 \times Z_2$ symmetry group

Klein symmetry: $\mathcal{K} = \{1, k_1, k_2, k_3\}$

for $U_{\text{PMNS}} = U_{\text{TB}}$:

$$k_1 = \frac{1}{3} \begin{pmatrix} -1 & 2 & 2 \\ 2 & -1 & 2 \\ 2 & 2 & -1 \end{pmatrix}, \quad k_2 = - \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \quad k_3 = k_1 k_2$$

Origin of the Klein symmetry $\mathcal{K}$

► direct models

- Klein symmetry $\mathcal{K} \subset$ family symmetry $\mathcal{G}$
- flavons ϕ are multiplets of $\mathcal{G}$
- their VEVs $\langle \phi \rangle$ break $\mathcal{G}$ down to $\mathcal{K}$ in neutrino sector
- for TBmax mixing (k_1, k_2, h) generate permutation group S_4

► indirect models

- Klein symmetry $\mathcal{K} \not\subset$ family symmetry $\mathcal{G}$
- $\mathcal{G}$ responsible for generating particular flavon VEV configurations $\langle \phi \rangle$
- for TBmax mixing – from e.g. $\Delta(27)$, $Z_7 \rtimes Z_3$

$$\langle \phi_1 \rangle \propto \begin{pmatrix} -2 \\ 1 \\ 1 \end{pmatrix} \quad \langle \phi_2 \rangle \propto \begin{pmatrix} 1 \\ 1 \\ 1 \end{pmatrix} \quad \langle \phi_3 \rangle \propto \begin{pmatrix} 0 \\ 1 \\ -1 \end{pmatrix}$$

$$\Rightarrow \quad \mathcal{L}_\nu \quad \sim \quad \nu (\phi_1 \phi_1^T + \phi_2 \phi_2^T + \phi_3 \phi_3^T) \nu \quad H \quad H$$

(exact) Tri-Bimaximal mixing is ruled out !

T2K [arXiv:1106.2822]

- $\theta_{13} \neq 0$ disfavored at $\sim 2.5\sigma$
- $5^\circ \lesssim \theta_{13} \lesssim 18^\circ$ at 90% C.L.

Double Chooz [arXiv:1207.6632]

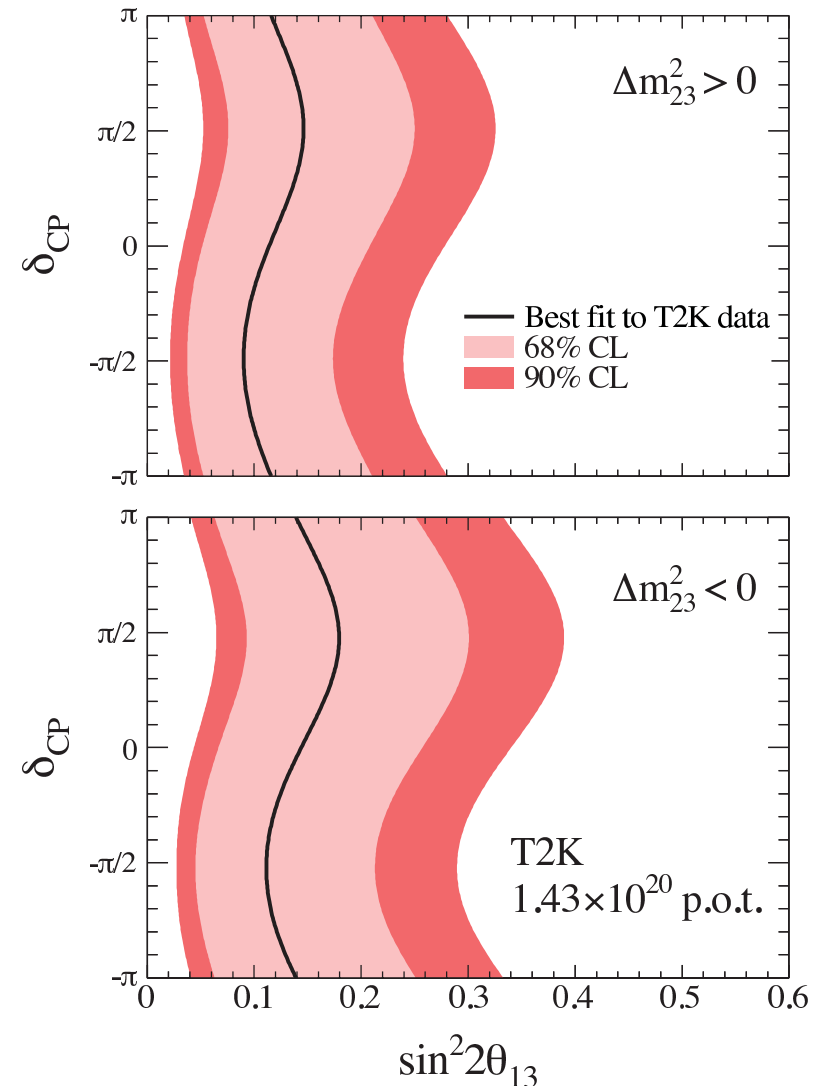
- $\theta_{13} \neq 0$ disfavored at $\sim 2.9\sigma$
- $6^\circ \lesssim \theta_{13} \lesssim 12^\circ$ at 90% C.L.

RENO [arXiv:1204.0626]

- $\theta_{13} \neq 0$ disfavored at $\sim 4.9\sigma$
- $8.0^\circ \lesssim \theta_{13} \lesssim 11.4^\circ$ at 90% C.L.

Daya Bay [arXiv:1210.6327]

- $\theta_{13} \neq 0$ disfavored at $\sim 7.7\sigma$
- $7.7^\circ \lesssim \theta_{13} \lesssim 9.6^\circ$ at 90% C.L.



Requires new model building strategies

direct models

indirect models

TBmax plus corrections

TBmax plus corrections

other family symmetries
with non-standard $\mathcal{K}$

non-standard flavon
VEV configurations

A direct model of leptons based on S_4

- discrete family symmetry with 24 elements
- irreducible S_4 representations: $\mathbf{1}$ $\mathbf{1}'$ $\mathbf{2}$ $\mathbf{3}$ $\mathbf{3}'$
- diagonal charged leptons
- in neutrino sector ($L \sim N^c \sim \mathbf{3}$ $H_u \sim \mathbf{1}$)

$$W_\nu \sim LN^c H_u + (\phi_{\mathbf{3}'} + \phi_{\mathbf{2}} + \phi_{\mathbf{1}})N^c N^c + \frac{1}{M} \tilde{\phi}_{\mathbf{1}'} \phi_{\mathbf{2}} N^c N^c$$

S_4 irrep	k_1	k_2	VEV alignment
$\mathbf{3}'$	$\frac{1}{3} \begin{pmatrix} -1 & 2 & 2 \\ 2 & -1 & 2 \\ 2 & 2 & -1 \end{pmatrix}$	$\begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$	$\langle \phi_{\mathbf{3}'} \rangle \propto \begin{pmatrix} 1 \\ 1 \\ 1 \end{pmatrix}$
$\mathbf{2}$	$\begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$	$\begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$	$\langle \phi_{\mathbf{2}} \rangle \propto \begin{pmatrix} 1 \\ 1 \end{pmatrix}$
$\mathbf{1}$	1	1	$\langle \phi_{\mathbf{1}} \rangle \propto 1$
$\mathbf{1}'$	1	-1	$\langle \tilde{\phi}_{\mathbf{1}'} \rangle \propto 1$

→ TBmax Klein symmetry broken in k_2 (at higher order)

Resulting neutrino mixing

$$U_{\text{PMNS}} \approx \begin{pmatrix} -\frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{2}{\sqrt{6}}\alpha^* \\ \frac{1}{\sqrt{6}} - \frac{1}{\sqrt{2}}\alpha & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} + \frac{1}{\sqrt{6}}\alpha^* \\ \frac{1}{\sqrt{6}} + \frac{1}{\sqrt{2}}\alpha & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}} + \frac{1}{\sqrt{6}}\alpha^* \end{pmatrix} \quad \begin{array}{l} \theta_{13} \approx 9^\circ \\ \text{requires} \\ |\alpha| \sim 0.2 \end{array}$$

- second column of $U_{\text{PMNS}} \propto (1, 1, 1)^T$
- trimaximal mixing – due to k_1 symmetry

-
- TBmax breaking caused by only one free (complex) parameter
 - affects θ_{13} and θ_{23} – but not θ_{12}
 - get correlations between mixing parameters

$\implies$ **testable** mixing sum rules

$$\theta_{23} \approx 45^\circ - \frac{1}{\sqrt{2}} \theta_{13} \cos \delta \qquad \theta_{12} \approx 35.3^\circ \text{ (TBmax value)}$$

Conclusion and outlook

- ▶ discrete symmetries essential in supersymmetry (proton decay)
 - $Z_n^{(R)}$ symmetries from $U(1)_{\text{FN}}^{(R)}$
 - R -parity violating symmetries might be interesting
- ▶ observed pattern in neutrino mixing
 - suggestive of non-Abelian (discrete) family symmetries
 - $\theta_{13} \approx 9^\circ$ from Daya Bay and RENO
 - deviations from tri-bimaximal mixing
 - testable mixing sum rules
- ▶ other aspects
 - ameliorate SUSY flavor problem
 - combine family and CP symmetry

Thank you

Neutrino oscillations: experimental milestones

► atmospheric neutrinos

- $\nu_\mu / \bar{\nu}_\mu$ disappear – Super-Kamiokande (1998)

► accelerator neutrinos

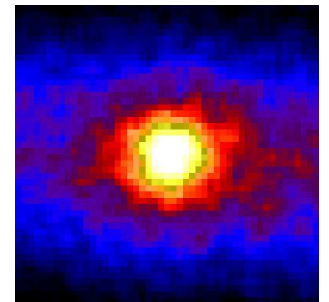
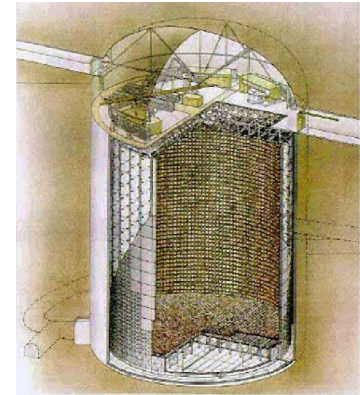
- ν_μ disappear – K2K (2002), MINOS (2006)
- ν_μ converted to ν_τ – OPERA (2010 & 2012)
- ν_μ converted to ν_e – T2K, MINOS (2011)

► solar neutrinos

- ν_e disappear – Chlorine (1998), Gallium (1999 - 2009),
Super-Kamiokande (2002), Borexino (2008)
- ν_e converted to $(\nu_\mu + \nu_\tau)$ – SNO (2002)

► reactor neutrinos

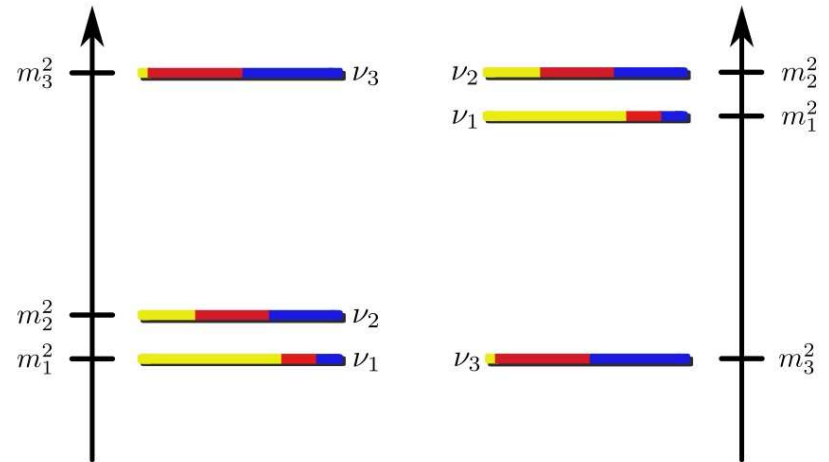
- $\bar{\nu}_e$ disappear – Double Chooz (2011), Daya Bay, RENO (2012)
- $\bar{\nu}_e$ disappear – KamLAND (2002)



Three neutrino flavor mixing

(in diagonal charged lepton basis)

$$\begin{array}{c} \text{flavor} \\ \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} \end{array} = \begin{array}{c} \text{PMNS mixing} \\ \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \end{array} \begin{array}{c} \text{mass} \\ \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \end{array}$$



$$U_{\text{PMNS}} = \begin{array}{c} \text{atmospheric} \\ \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \end{array} \begin{array}{c} \text{reactor + Dirac} \\ \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \end{array} \begin{array}{c} \text{solar} \\ \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \end{array} \begin{array}{c} \text{Majorana} \\ \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{\frac{\alpha_2}{2}} & 0 \\ 0 & 0 & e^{\frac{\alpha_3}{2}} \end{pmatrix} \end{array}$$

$$\theta_{23} \approx 45^\circ \quad \theta_{13} \approx 9^\circ \quad \theta_{23} \approx 33^\circ$$