GR 4: Klassische Allgemeine Relativitätstheorie I

Zeit: Dienstag 15:40-16:20

 $\begin{array}{cccccc} & & GR \ 4.1 & Di \ 15:40 & 30.45: \ 101 \\ \textbf{Linearized gravity on type D backgrounds} & & Steffen \\ & & AKSTEINER^{1,2} \ and \bullet LARS \ ANDERSSON^3 & & ^1Quest \ Uni \ Hannover, \\ & Deutschland & & ^2Zarm \ Uni \ Bremen, \ Deutschland & & ^3MPI, \ Golm, \\ & Deutschland \end{array}$

In this talk I present joint work with Lars Andersson about the field equations of linearized gravity on a Petrov type D background, which includes Kerr spacetime. The Geroch Held Penrose (GHP) formalism is used to derive decoupled equations for all linearized Weyl scalars. The identification of gauge source functions leads to a generalized Regge-Wheeler equation. On Schwarzschild, a derivation of the gauge invariant Regge-Wheeler and Zerilli equation directly from the equation for the spin 0 scalar will be presented.

 $GR \ 4.2 \quad Di \ 16:00 \quad 30.45: \ 101 \\ \textbf{Properties of the 1-PN Dedekind ellipsoids} \ - \ \bullet \texttt{Norman}$

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A changing quadrupole moment leads to gravitational radiation in General Relativity. Does this imply that stationary but nonaxisymmetric, isolated systems cannot exist? To learn something about the answer to this question, a post-Newtonian (PN) approximation of the Newtonian triaxial and homogeneous Dedekind ellipsoids is investigated. We shall discuss a generalization of the ansatz used by Chandrasekhar and Elbert (1978), in particular its axisymmetric limit. Contrary to Chandrasekhar & Elbert's ansatz this generalization permits an axially symmetric and rigidly rotating limit (PN Maclaurin spheroids). The additional freedom in the generalized solution can also be used to remove a singularity which occurs in their work. A limit where the Dedekind ellipsoids degenerate to a line mass distribution is also discussed.